



# **Astrosat and Chandrayaan-1 Some Evolutionary Considerations**

**K Kasturirangan**

**Foundation Day Function  
Raja Ramanna Centre for Advanced Technology, Indore  
19 February 2008**



The ROHINI satellite

## Rohini Satellite



Vikram Sarabhai



## Homi Bhabha Flying a Balloon



## Hard X-ray Detector: Balloon Payload

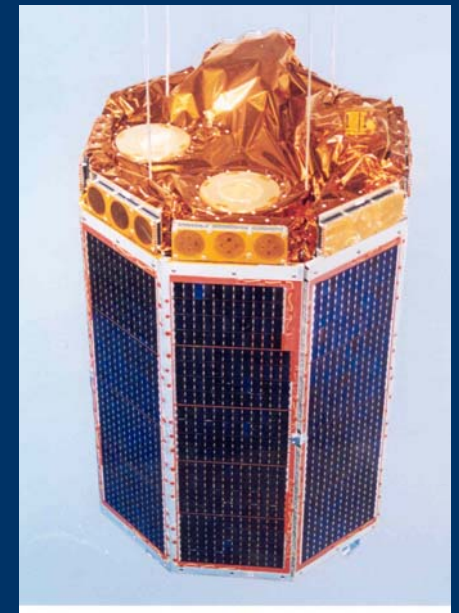
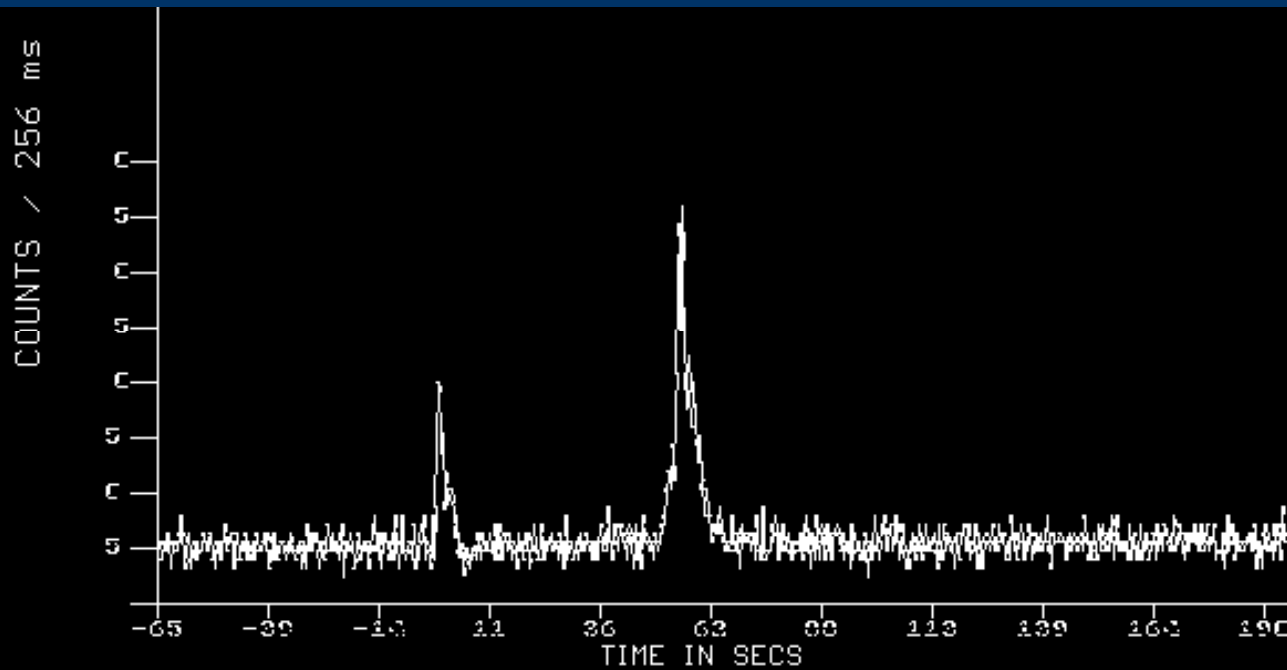


## Rohini Sounding Rockets



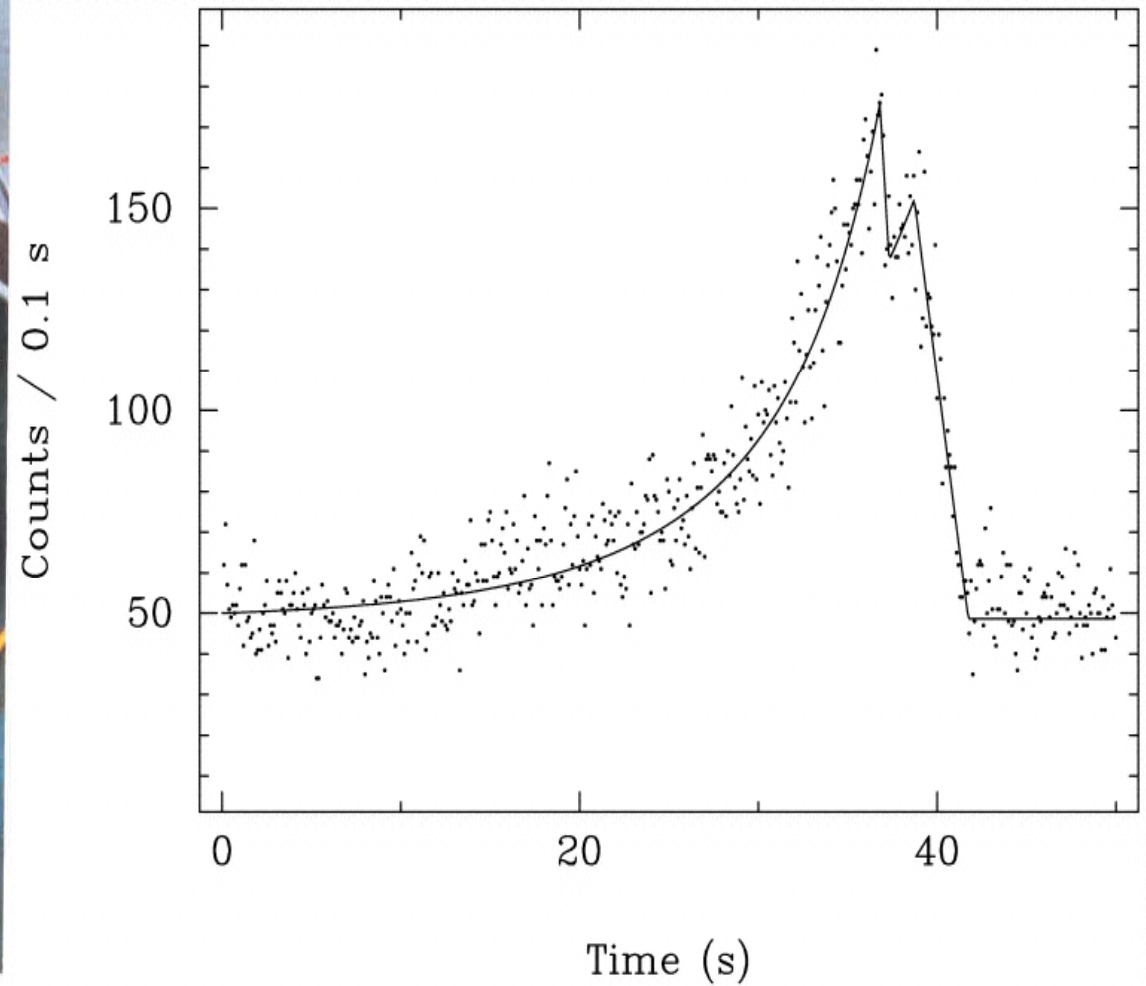
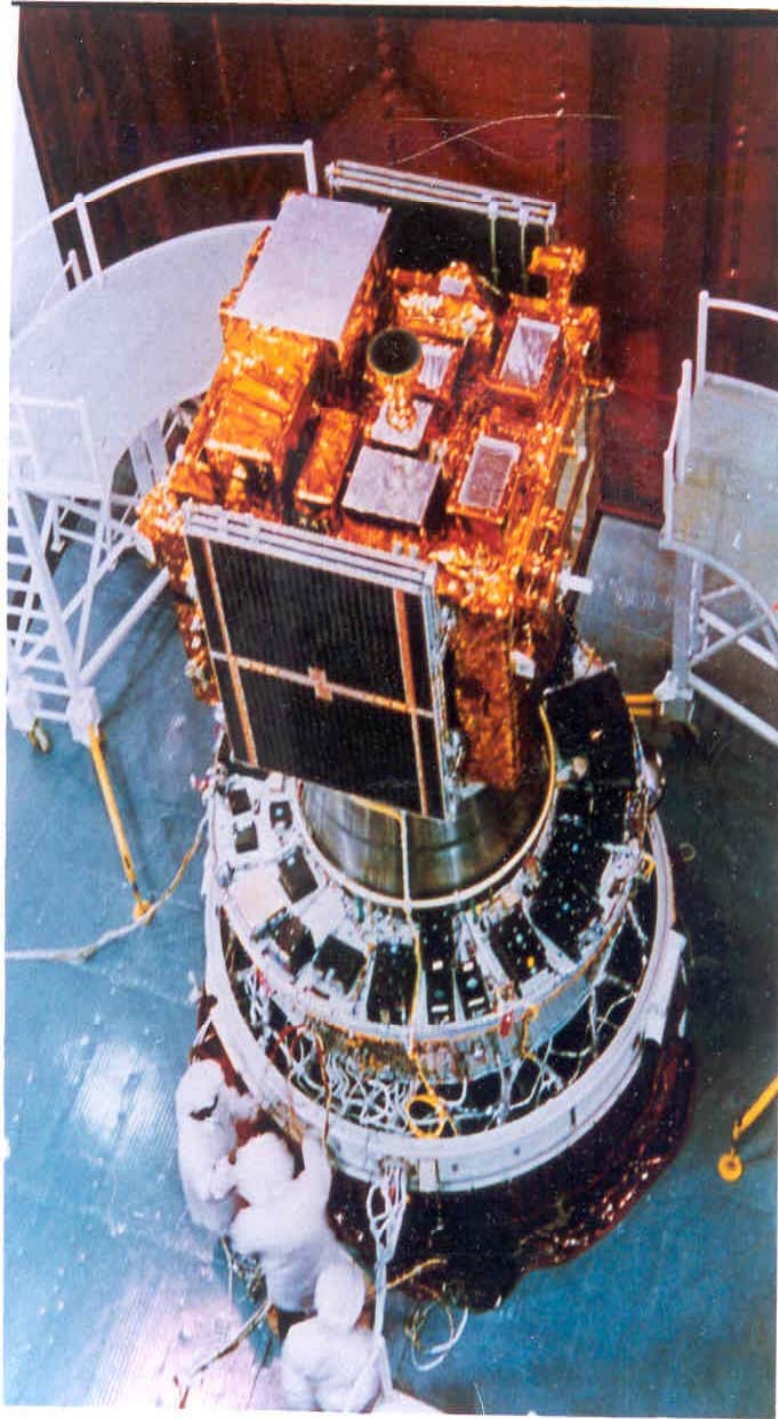
## Balloon Launch at Hyderabad

# Gamma Ray Burst experiment on SROSS (1994)





# IXAE on IRS-P3 (1996)





# Indian Remote Sensing Satellites

India has established the largest constellation of remote sensing satellites providing data in a variety of spatial resolutions and spectral bands for different applications.

**Mission Completed**

**Currently in Operation**



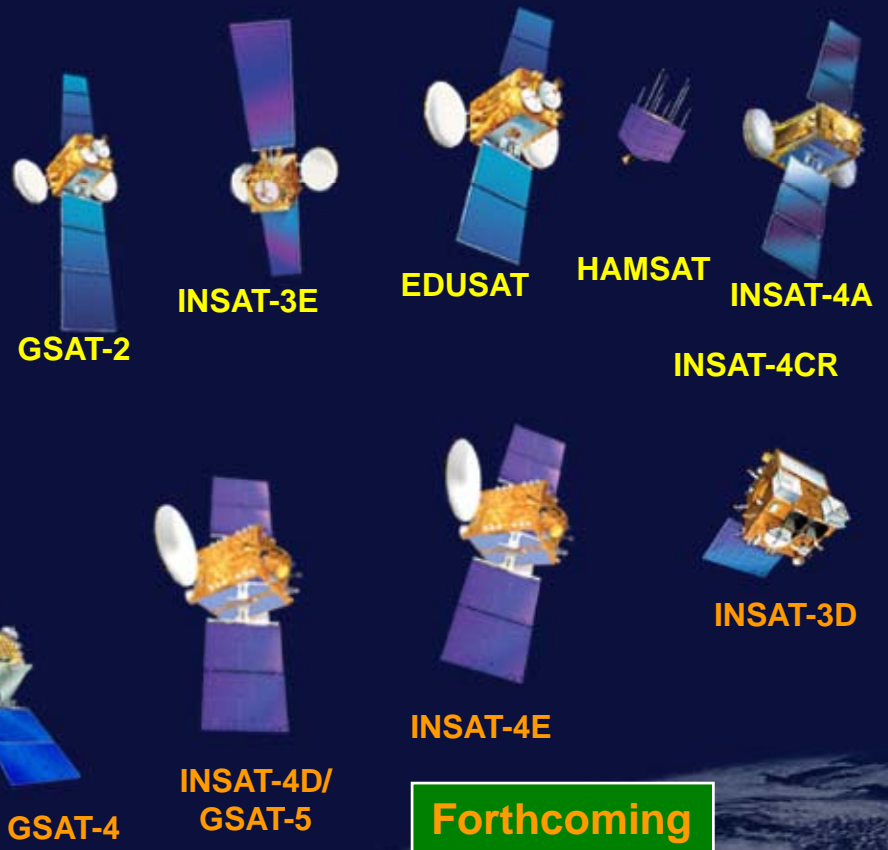
# INSAT Satellites

Indian National Satellite (INSAT) System established in 1983, is a multipurpose system for telecommunications, television broadcasting and radio networking, meteorology and disaster warning.

Mission Completed



Currently in Operation



Forthcoming





# ISRO LAUNCHERS



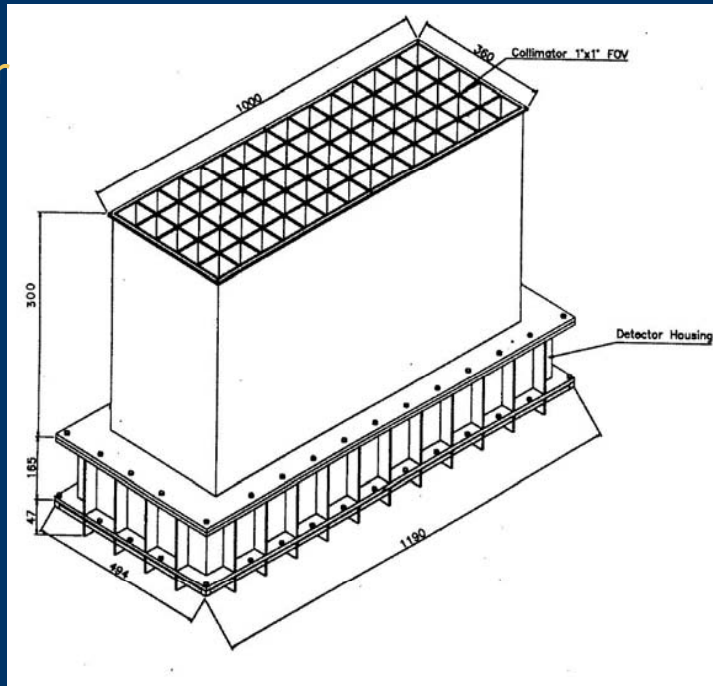
	PSLV	GSLV	GSLV MkIII
Weight (T)	294	400	629
Payload (Kgs)	1500 SSO	2250 GTO	4000 to 4500 GTO
Flights	11 (1993-07)	5 (2001-07)	--

# Scientific objectives of ASTROSAT

- To understand the broad-band x-ray emission spectrum in cosmic sources.
- To study correlated intensity variations over time in the visible, UV, soft and hard X-ray bands to address the origin of radiation in the different wave bands.
- Search for black hole sources by limited surveys in the galactic plane.
- Measure magnetic fields of neutron stars by detection and studies of cyclotron lines in the X-ray spectra of X-ray pulsars.
- Detect and locate new transient X-ray sources.
- Multi-band survey covering Ultra-violet band from 130-300 nm and X-ray band from 0.3 - 100 keV .
- Deep surveys of selected regions of the sky to detect faint quasars to study their clustering and large scale structures, and obtain UV fluxes from very distant galaxies.

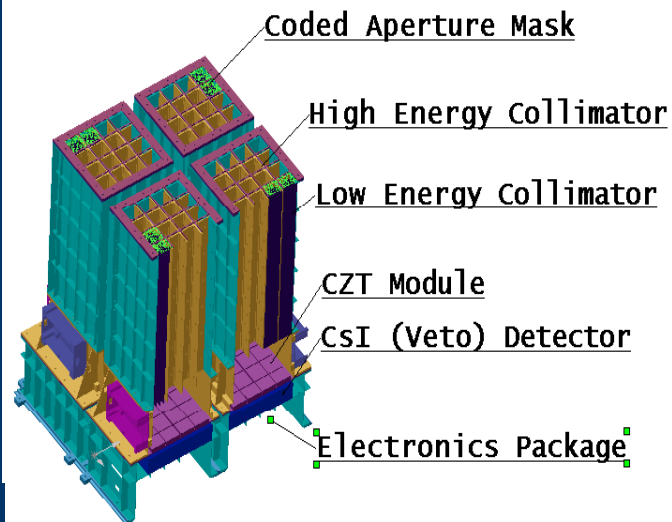


## Large Area Xenon Proportional Counter (LAXPC)



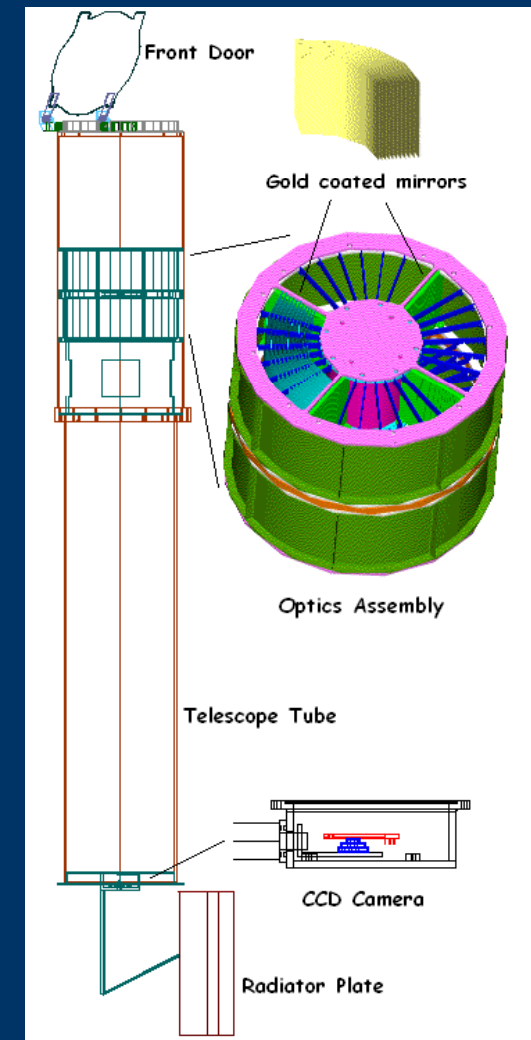
Energy Range 2 – 80 keV  
Effective area ~ 6000 cm<sup>2</sup>

## CZT imager

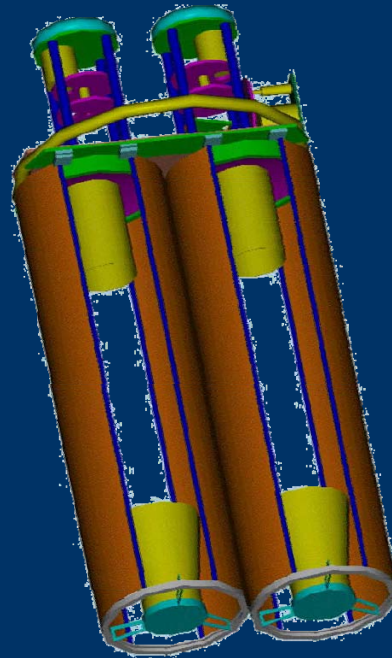
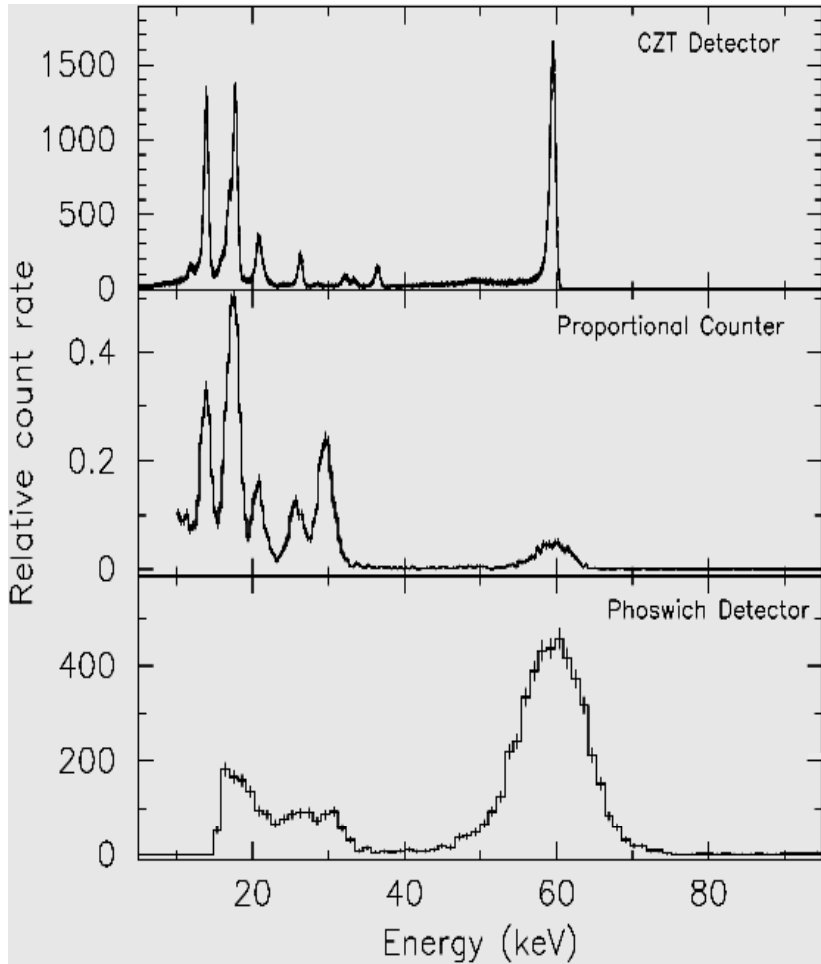


Energy Range 10 – 100 keV  
Resolution = 5% @ 60 keV  
Effective area ~ 1000 cm<sup>2</sup>

Energy Range 0.5 – 8 keV  
Focal length = 2 m  
Effective area ~ 200 cm<sup>2</sup>



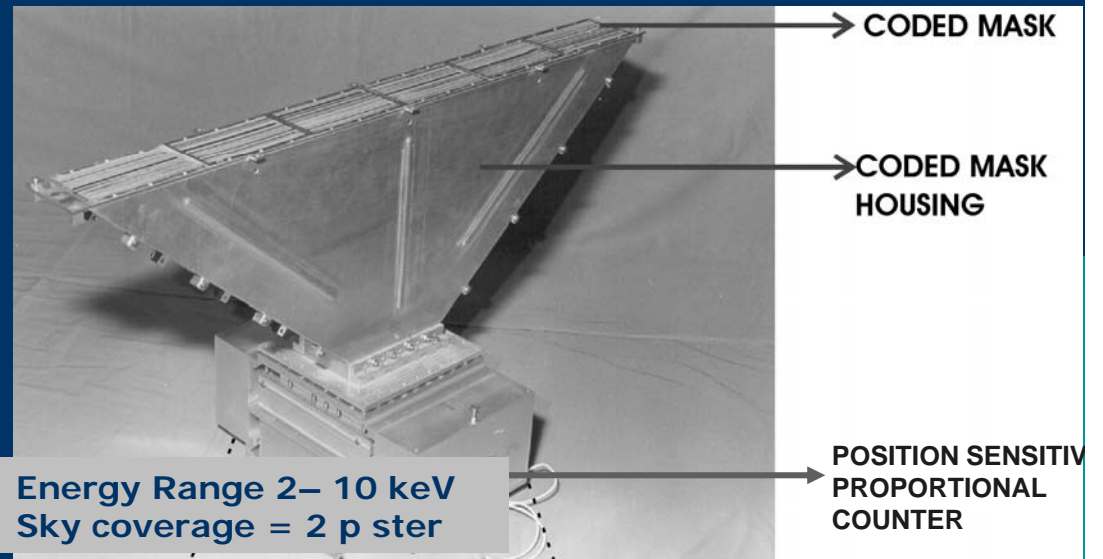
## Soft X-ray telescope (SXT)



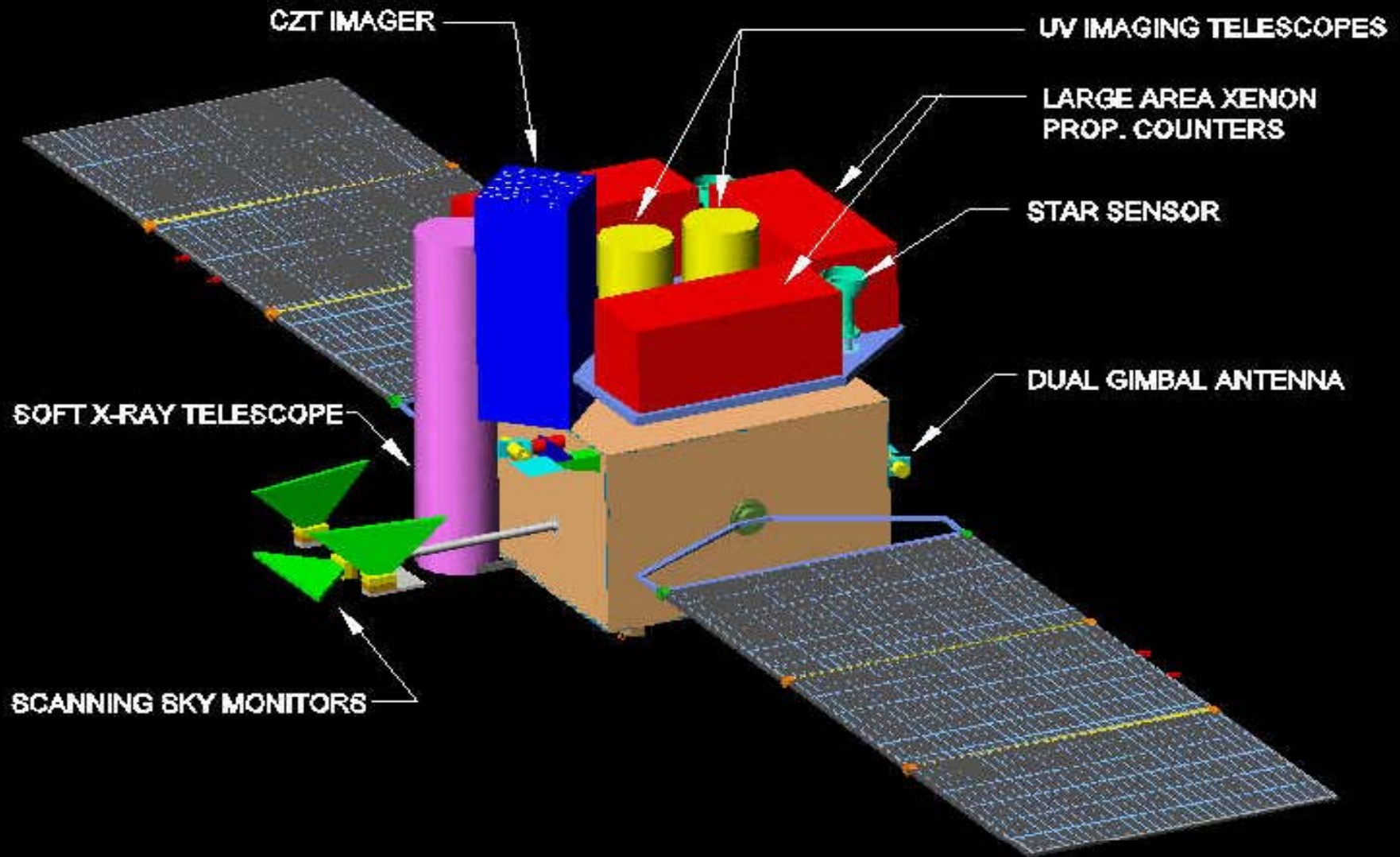
## UV Imaging Telescope (UVIT)

Visible (3500-6000 Å)  
 Near UV (1800-3000 Å)  
 Far UV (1300-1800 Å)  
 Ang. res = 1.8 arc sec

## Scanning Sky Monitor (SSM)



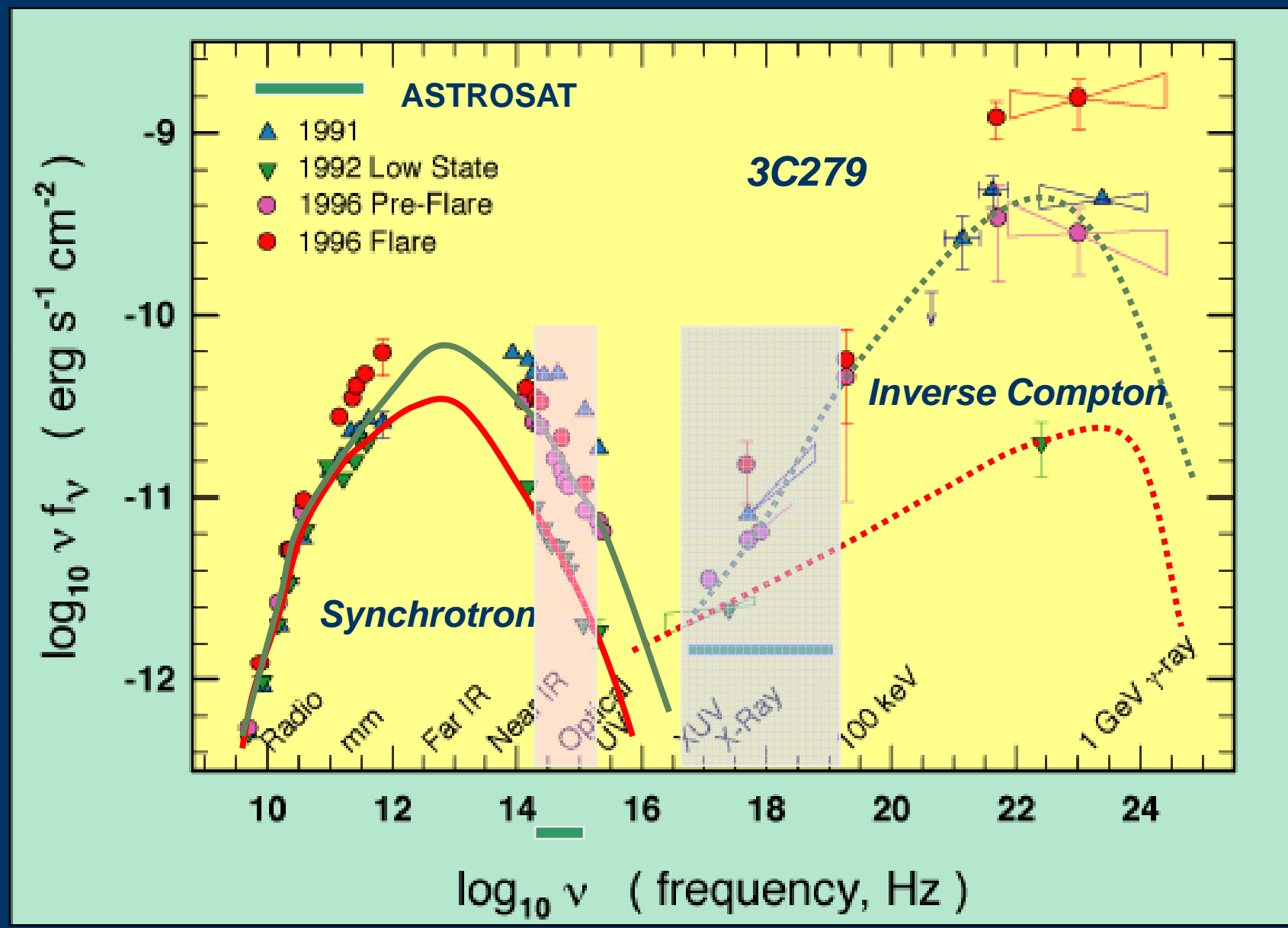
# ASTROSAT



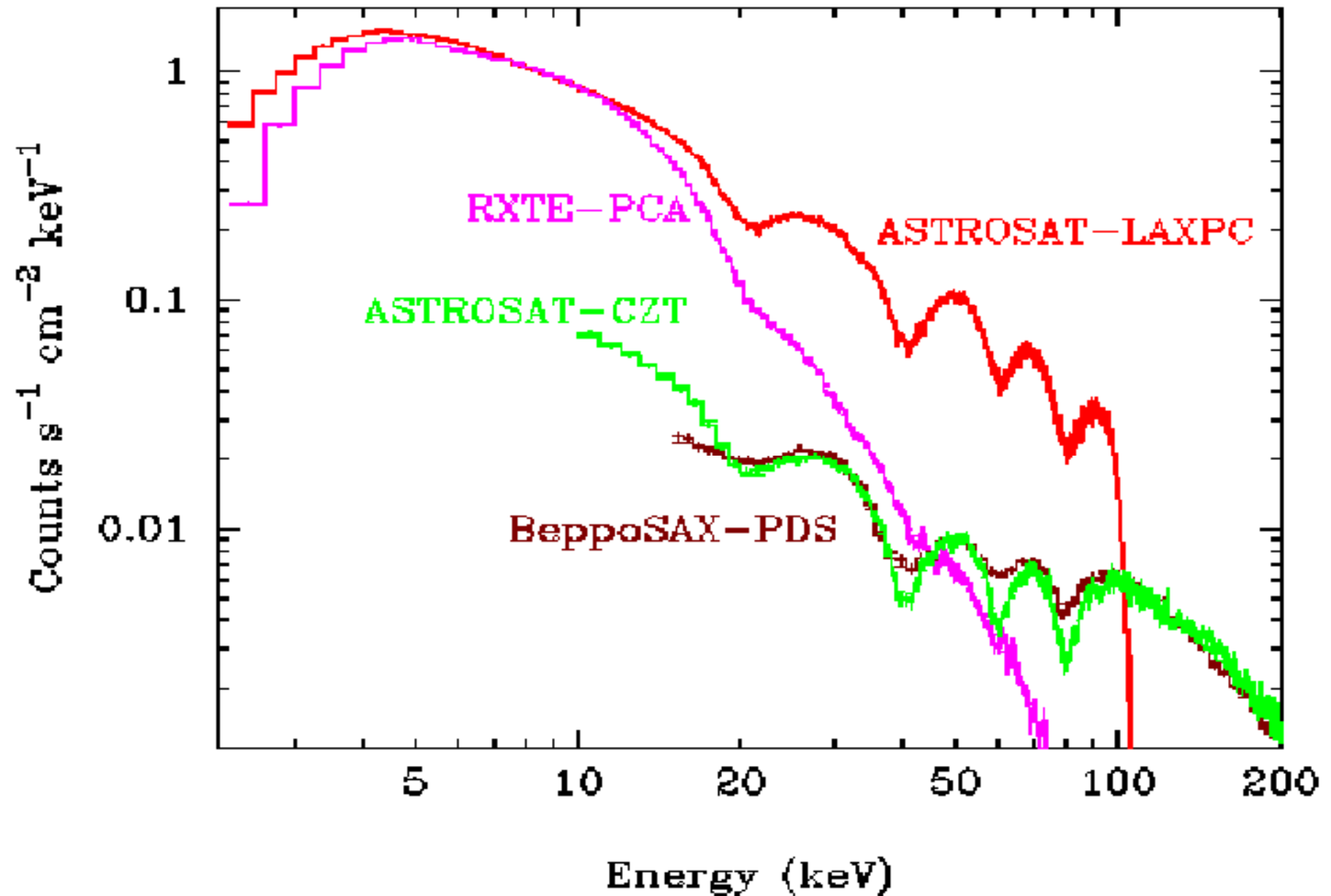
**DEPLOYED CONFIGURATION OF ASTROSAT**



# Simultaneous Multiwavelength Coverage: Spectral



# Detection of Cyclotron Lines



# Early work on Lunar Science

## Remote sensing

- Resourcesat
- Cartosat (2.5 m stereoscopic resolution)





# CHANDRAYAAN-1: India's Mission to the Moon

- ◆ Will address some of the outstanding questions concerning the origin and evolution of the Moon.
- ◆ Will give impetus to planetary research through a committed long-term program.
- ◆ Will provide unique opportunities to upgrade several areas of technology, and
- ◆ Will help to create a cadre of young scientists who can also participate in future international planetary missions.

# Theories

- **Fission**
- **Capture**
- **Accretion**
- **Giant impact**



Origin?

# Understanding the origin and Evolution of the Moon

## Physical Properties of the Moon

Topography

Gravity

Magnetic Field

Radiation Environment



## The bulk chemistry of Moon

Nature of the Lunar Crust

The Lunar Far-side:  
Rock types, Chemistry

### *Special Regions of Interest:*

*Polar Regions ,*

*South Pole Aitken Region,*

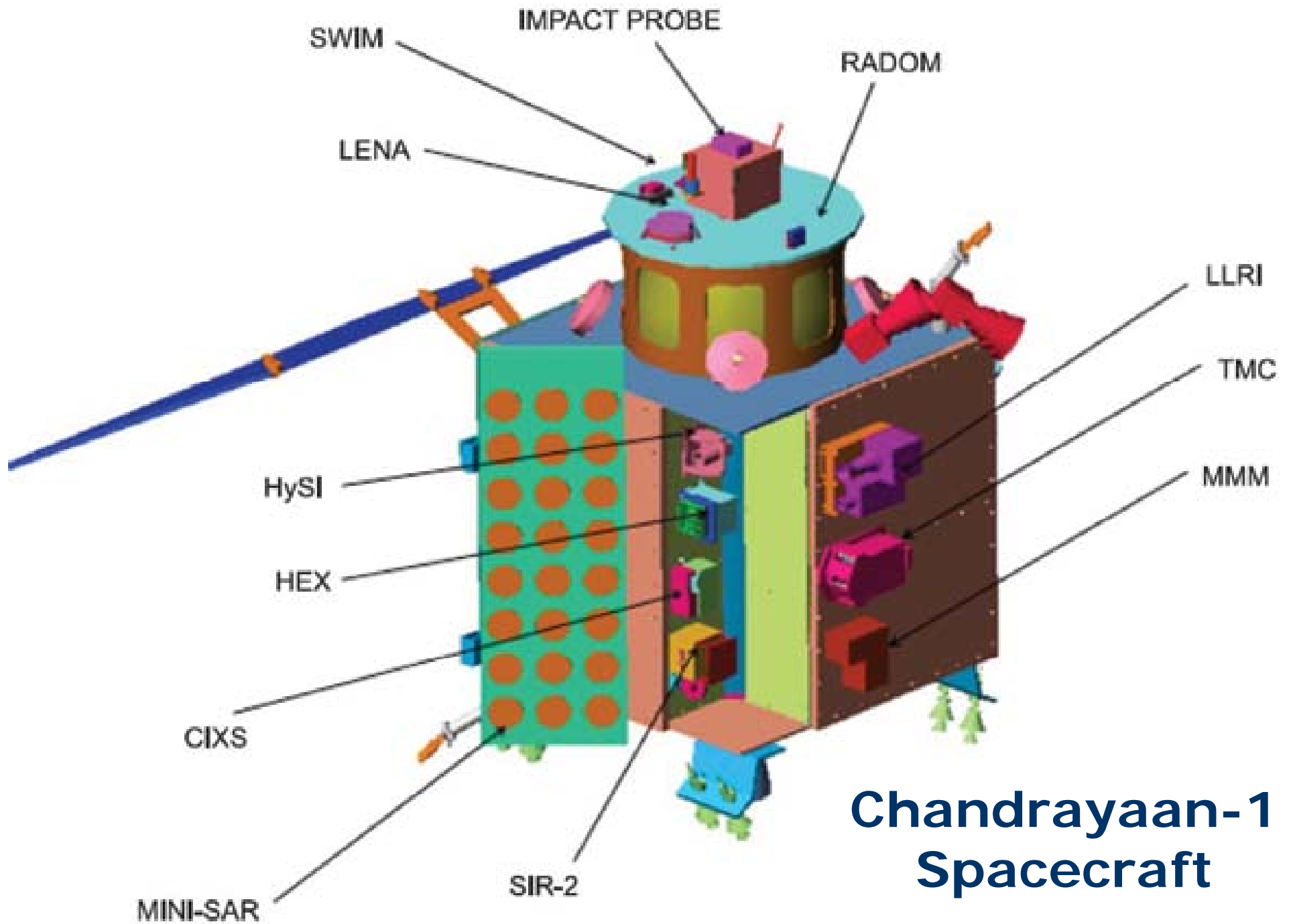
*Selected Basins and Craters with central uplift*

*Nature of Volatile Transport on Moon (Water on Moon?)*



## Chandrayaan-1 payloads

Payload	Sensor Configuration	Wavelength/energy Range	Spatial Resolution	Objective
Hyperspectral imager (HySI)	Wedge filter pixelated imager	0.4–0.92 $\mu\text{m}$ with 15 nm resolution using 64 channels	80 m	Areal mapping of minerals
Infra red spectrometer (SIR-2)	Grating spectrometer	0.93–2.4 $\mu\text{m}$	100 m	Linear mapping Of minerals
Moon Mineral Mapper (M <sup>3</sup> )	Grating spectrometer and HgCdTe detector	0.7 to 3.0 $\mu\text{m}$ with 10 nm resolution	30 m	Areal mineral and resource (water, organics) mapping
Terrain mapping Camera (TMC)	Three stereo cameras with pixelated detectors	Panchromatic	10 m areal 5 m elevation	Topographic mapping
Laser Ranging (LLRI)	Pulsed Nd-Yag laser with optical system	1064nm	Elevation 10 m	Topography, Chandrayaan altimetry
X-ray Fluorescence spectrometer (C1XS)	Swept charged CCD	1–10 keV	20 km	Chemical mapping (Mg, Al, Si, Ca, Ti, Fe)
Solar X-ray Monitor (XSM)	Si pin diode	2–10 keV	–	Solar X-ray spectrum
High energy X-ray spectrometer (HEX)	CdZnTe detector	20–250 keV	40 km	Th, <sup>210</sup> Pb mapping
Synthetic Aperture Radar (mini SAR)	Radar, Scatterometer and altimeter	2.4 GHz	100 m	Soil properties Topography, altimetry
Neutral atom analyzer (SARA)	Mass spectrometer and solar wind monitor	10eV-keV	100 m	atmospheric neutrals (H-Fe) composition, Magnetic anomalies,
Radiation Dose Monitor (RADOM)	Si semiconductor	>8keV	–	Radiation dose



# Chandrayaan-1 Mission Phase

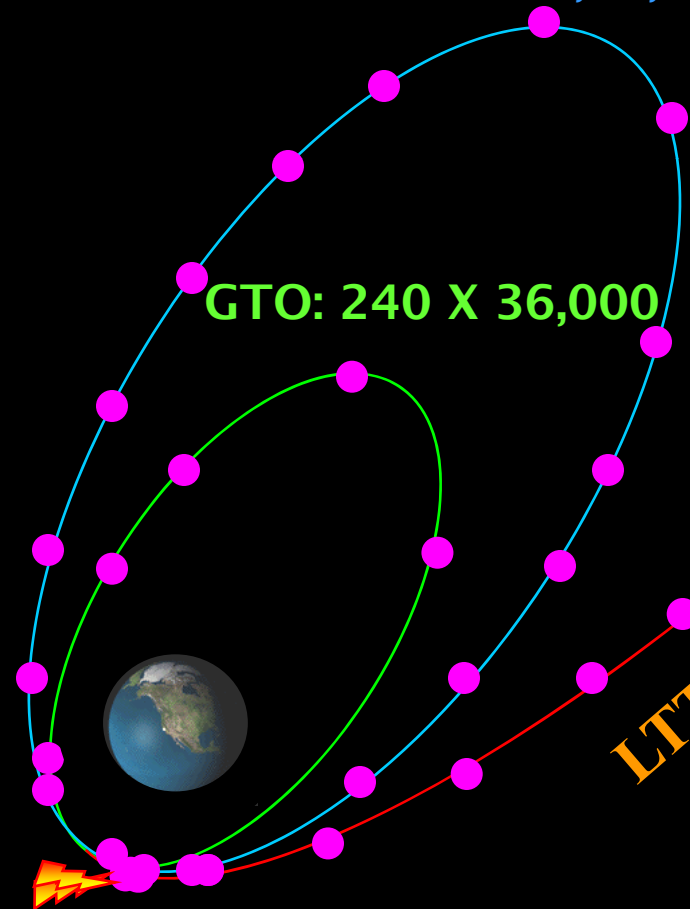


SUN

ETO: 240 X 1,00,000 km

GTO: 240 X 36,000

LTT: 240 X 3,86,000 Km



PERIGEE MANEUVERS

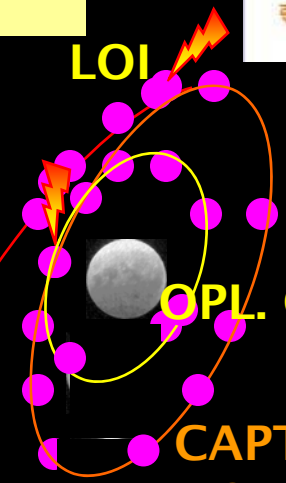
MCC

MCC

LOI

OPL. ORBIT

CAPTURE ORBIT



PROGRESS OF MOON IN ITS ORBIT



"No Country which wishes to play a leading part in the world can afford to neglect pure or long term research"

Homi J Bhabha





Thankyou